Abstract—Atmospheric propagation models and the measurements that train them are critical to the design of efficient and effective space-ground links. As communication systems advance to higher frequencies in search of higher data rates and open spectrum, a lack of data at these frequencies necessitates new measurements to properly develop, validate, and refine the models used for link budgeting and system design. In collaboration with the Air Force Research Laboratory (AFRL), NASA Glenn Research Center has deployed the W/V-band Terrestrial Link Experiment (WTLE) in Albuquerque, NM to conduct a measurement campaign at 72 and 84 GHz, among the first atmospheric propagation measurements at these frequencies. WTLE has been operational since October 1, 2015, and the system design shall be herein discussed alongside preliminary results and performance.

Keywords—millimeter wave propagation, propagation losses, radiowave propagation.

I. INTRODUCTION

While NASA’s communications architectures are currently moving toward Ka and optical frequencies, the agency is also investigating available spectrum in the Q/V/W-bands as a downlink option for the next generation Space Based Relay (SBR) to superecede the Tracking and Data Relay Satellite (TDRS) system in coming years [1]. Accordingly, NASA Glenn Research Center (GRC) and the Air Force Research Laboratory (AFRL) have initiated a joint terrestrial propagation campaign to characterize rain attenuation, depolarization, scintillation, and gaseous absorption effects of the atmosphere in the V and W-bands (72 GHz and 84 GHz). The W/V-Band Terrestrial Link Experiment (WTLE) consists of a transmitter on the crest of the Sandia Mountains and a receiver on the roof of a University of New Mexico research center (COSMIAC) in south Albuquerque. The slant path is approximately 24 km long, with the receiver and transmitter at elevations of 1.6 km and 3.2 km respectively, resulting in a look angle of 3.9° above the horizon. Fig. 1 shows the location of the transmitter at Sandia Crest, the receivers on the roof of the COSMIAC building, and an elevation profile of the slant path covered by the link. The link was installed in late September 2015 and has been operational since October 1, 2015.

II. SYSTEM DESIGN

The WTLE transmitter consists of a coherent 72 and 84 GHz continuous wave (CW) beacon with an EIRP of 40 dBm using two lens antennas with a 3° half-power beam-width and 35 dBi directivity. The transmitter, including the lens antennas, is contained in a weather-proof enclosure with the electronics temperature controlled to within ±0.01°C and air temperature controlled to within ±0.1°C. Polarization is LHCP from the perspective of the receivers.

The receiver system consists of two 0.5m V and W-band Cassegrain reflectors. The twin receivers observe both the co- and cross-polarization components of each channel and downconvert the signals to 7 MHz at the feed before digitization. The receiver electronics are also temperature controlled to within ±0.01°C, and a calibration tone is injected before the LNA for use in periodically monitoring overall system gain. Dynamic range, isolation, and other specifications of the receiver are summarized in Table 1. After digitization, the signals are processed using a Fast Fourier Transform (FFT) I/Q algorithm utilized in previous NASA propagation terminals [2] which uses a frequency estimation technique to coherently track and measure the amplitude of the beacon signals [3, 4]. An overall measurement rate of 10 Hz rate is implemented to characterize scintillation effects. The receive site also has a variety of colocated meteorology equipment including a weather station, laser disdrometer, and SODAR. Weather instrumentation is also planned for installation at the transmit site and at an intermediate location along the path. More information on the experiment design and instrumentation is presented in [5].
TABLE I. RECEIVER SPECIFICATIONS

<table>
<thead>
<tr>
<th>Channel</th>
<th>Parameter</th>
<th>Spec.</th>
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</thead>
<tbody>
<tr>
<td>V (72 GHz)</td>
<td>Dynamic Range</td>
<td>70 dB</td>
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<tr>
<td></td>
<td>Co/Cross-Polarization Isolation</td>
<td>13 dB</td>
</tr>
<tr>
<td>W (84 GHz)</td>
<td>Dynamic Range</td>
<td>68 dB</td>
</tr>
<tr>
<td></td>
<td>Co/Cross-Polarization Isolation</td>
<td>20 dB</td>
</tr>
<tr>
<td></td>
<td>Polarization</td>
<td>LHCP</td>
</tr>
<tr>
<td></td>
<td>Measurement Rate</td>
<td>10 Hz</td>
</tr>
</tbody>
</table>

III. PRELIMINARY RESULTS

The link has not yet been operational long enough to analyze long term results with statistical significance, however, several notable examples of various weather phenomena have been observed and are presented here as preliminary findings to demonstrate system performance. Figs. 3 – 5 plot the time series power measurements of the co- and cross-polarization channels at both V (72 GHz) and W (84 GHz) for various weather events. During clear sky conditions, the received power levels are approximately -14 dBm (co) and -27 dBm (cross) for the V-band, and -17 dBm (co) and -37 dBm (cross) for the W-band.

Firstly, an example of a deep rain fade event is presented in Fig. 3. For approximately 24 minutes during the event, the measured attenuation exceeds the system dynamic range of 70 dB on all channels. The concurrent rain rate measured at the receiver site peaked at 17.2 mm/hr, during this time, but may have been stronger along the path given the severe attenuation. No significant depolarization was observed throughout the event. However, in Fig. 4, a snow event is presented in which strong depolarization effects were measured. In fact, depolarization was observed to the extent that the W-band cross-polarization power level exceeded its respective co-polarization power during the height of the event. This event corresponds to a cumulative snowfall of 7.6 mm in Albuquerque throughout the day. Lastly, Fig. 5 presents an example of observed cloud attenuation over a period of 6 hours. The WTLE system is uniquely situated to measure cloud effects given that clouds often form along the path. As clouds pass along the link throughout the event in Fig. 5, attenuation rapidly fluctuates by up to 12 dB (V) and 15 dB (W), while the isolation between the co- and cross-polarizations remains consistent on both channels.

IV. CONCLUSIONS

Herein, we presented preliminary results from the joint NASA-AFRL terrestrial link experiment at W/V-band, which has been operating in Albuquerque since October 2015. The system has thus far performed as expected and recorded reliable data, including several examples of deep rain fades, depolarization effects, and cloud attenuation, which will ultimately be used to validate theoretical models of atmospheric propagation phenomena at V and W band. Future work for this campaign include the installation of additional weather instrumentation at the transmitter site and along the slant path, as well as a more thorough analysis of the collected data once the system has been operational for a statistically significant length of time.

REFERENCES